

## Homework 8

### P 2.1 – counts as 2 problems

### M 10-4 see below

[ $\Lambda$  is the thermal de Broglie wavelength,  $p$  is the pressure.]

### M 10-27 see next page

[10-26 is only for background.]

### M 10-29 see next page

[Interpretation for us: Get  $E$  from our grand partition function for the radiation field, then use this and energy flux =  $(1/4)(E/V)c$  to get the Stefan Boltzmann law, evaluating  $\sigma_B$ .]

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**10-4.** Derive Eqs. (10-11) and (10-12) from Eqs. (10-9) and (10-10).

$$N = 2\pi \left(\frac{2m}{h^2}\right)^{3/2} V \int_0^\infty \frac{\lambda \varepsilon^{1/2} e^{-\beta \varepsilon} d\varepsilon}{1 + \lambda e^{-\beta \varepsilon}} \quad (10-9)$$

$$pV = 2\pi kT \left(\frac{2m}{h^2}\right)^{3/2} V \int_0^\infty \varepsilon^{1/2} \ln(1 + \lambda e^{-\beta \varepsilon}) d\varepsilon \quad (10-10)$$

$$\rho = \frac{1}{\Lambda^3} \sum_{l=1}^{\infty} \frac{(-1)^{l+1} \lambda^l}{l^{3/2}} \quad (10-11)$$

$$\frac{p}{kT} = \frac{1}{\Lambda^3} \sum_{l=1}^{\infty} \frac{(-1)^{l+1} \lambda^l}{l^{5/2}} \quad (10-12)$$

[next page]

**10-26.** Show that the Planck blackbody distribution can be written in terms of wavelengths  $\lambda$  rather than frequency:

$$\rho(\lambda, T) d\lambda = \frac{8\pi hc}{\lambda^5} \frac{d\lambda}{e^{hc/\lambda kT} - 1}$$

where  $\rho(\lambda, T) d\lambda$  is the amount of energy between wavelength  $\lambda$  and  $\lambda + d\lambda$ .

**10-27.** If  $\omega_{\max}$  is the frequency at which  $\rho(\omega, T)$  is a maximum, illustrate by maximizing  $\ln \rho(\omega, T)$  that  $\omega_{\max}$  is given by

$$\frac{\hbar\omega_{\max}}{kT} = 3(1 - e^{-\hbar\omega_{\max}/kT})$$

and so

$$\frac{\hbar\omega_{\max}}{kT} = 2.82$$

Similarly show that

$$\lambda_{\max} T = 0.290 \text{ cm-deg}$$

**10-29.** In Problem 7-24, it was shown that the number of molecules striking a surface per unit area per unit time is  $\rho\bar{v}/4$ . By a similar approach, show that the total energy flux radiated by a blackbody is

$$e(T) = \frac{c}{4} \frac{E}{V} = \sigma T^4$$

where  $\sigma = 2\pi^5 k^4 / 15h^3 c^3$ . This result is known as the Stefan-Boltzmann law, and  $\sigma$  is the Stefan-Boltzmann constant. Verify that  $\sigma$ , a universal constant, equals  $5.669 \times 10^{-5} \text{ erg/cm}^2 \text{-deg}^4 \text{-sec}$ .